

Reading: Chap. 11, Chap. 4, Sect. 4.6, Chap.12, Sect. 12.1-12.5

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Homework #7: Due in recitation Friday/Monday (Oct. 25/28)

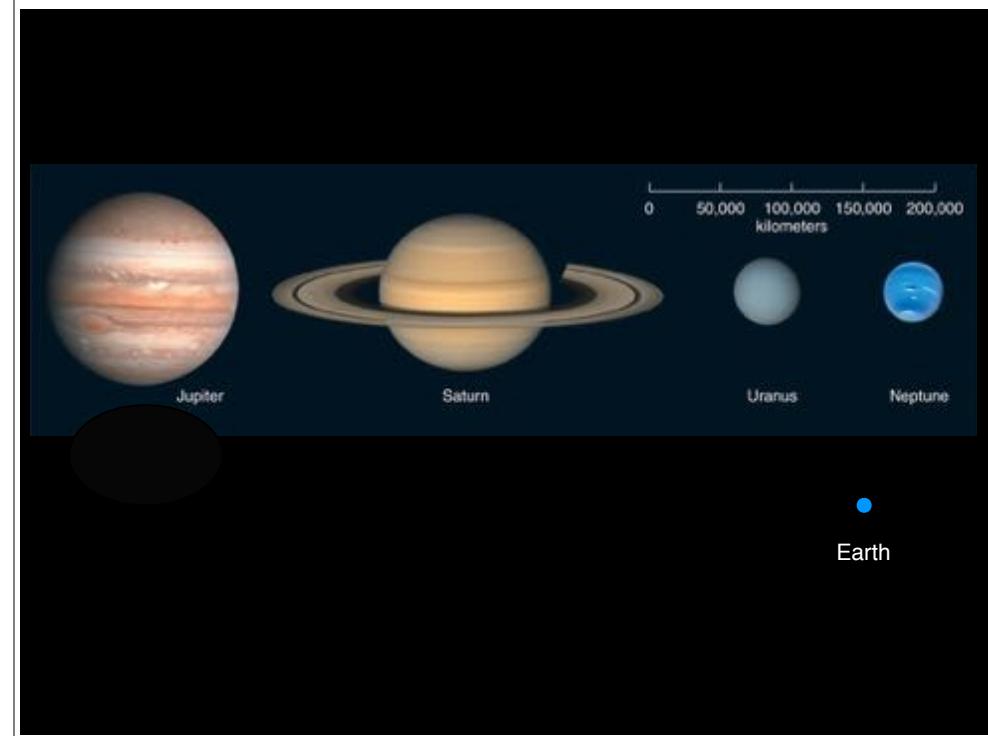
Homework #8: On website, due in recitation on Nov. 1/4

Last time: Planetary Atmospheres

- [Survey of Planetary Atmospheres](#)
- [Primary Atmosphere, Secondary Atmosphere alteration](#)
- [Atmospheric Pressure and Temperature](#)
 - Pressure vs. height: Hydrostatic Equilibrium; T vs. height: thermal equilibrium
- [Earth's Atmosphere](#)
- [The Greenhouse Effect](#)

Today: The Gas and Ice Giant Planets

- Jupiter (and Saturn)
 - Atmosphere Composition (H, He) and cloud coloration
 - Zones, Bands and circulation patterns
 - Interior
- Differences between Jupiter and Saturn
- Uranus and Neptune

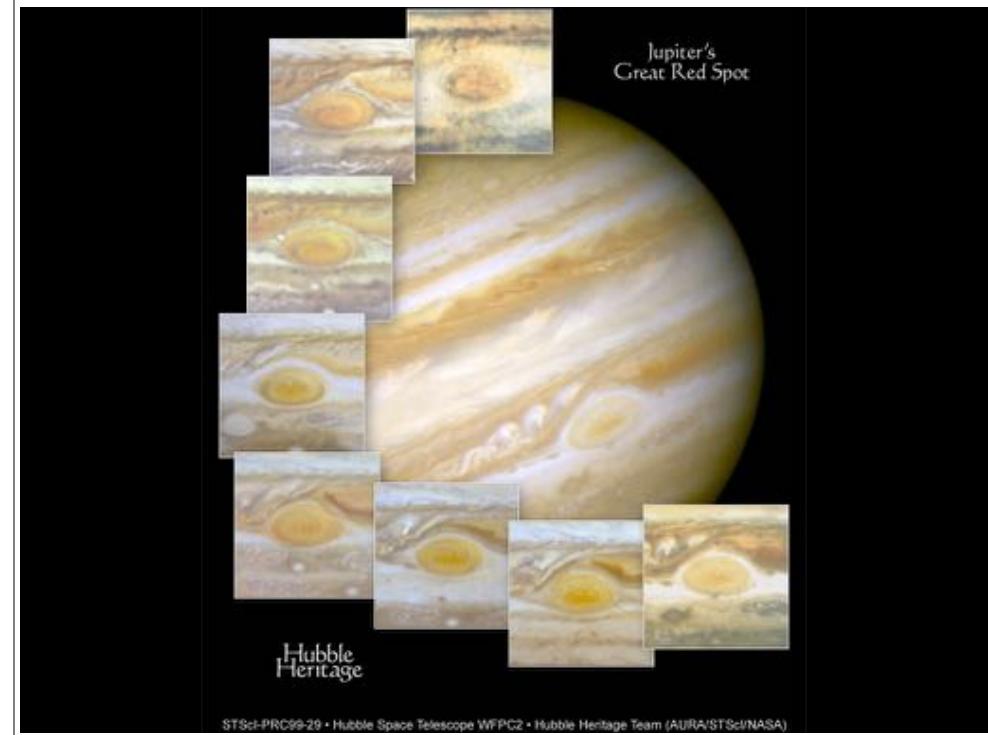


Jupiter

• The Basics:

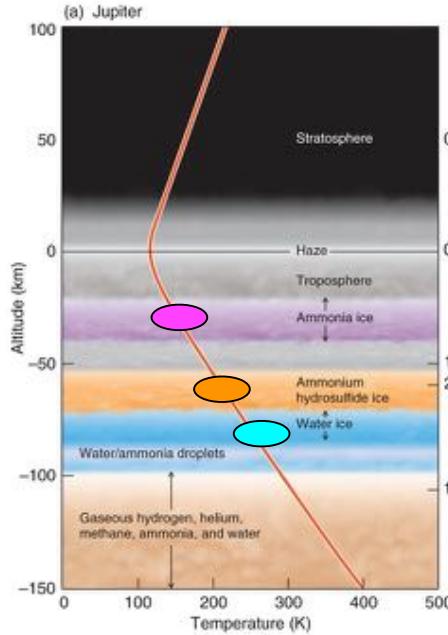
- Mass = 318 x Earth
- Diameter = 11.2 x Earth
- "Surface" Gravity = 2.53 x Earth
- Moons: 4 major, many minor
- Rotation: very fast: $P \sim 10$ hours
- view from Earth: cloud belts, no surface
- visits:
 - flybys: Pioneer 10, 11 (1973), Voyager 1, 2 (1979)
 - orbiter: Galileo (1995-2003), Juno (2016- 2018+)
 - atmosphere probe: Galileo (1995)

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The Jovian Atmosphere

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- 75% Hydrogen, ~25% Helium
- trace compounds (all H rich):
 - ammonia (NH_3)
 - methane (CH_4)
 - water vapor
- composition of clouds:
 - ammonia (high level)
 - ammonium hydrosulfide ice (middle level)
 - water ice (low)
- colors of clouds: only small amt. needed.
 - high white clouds (ammonia)
 - lower clouds are dark (amm. hydrosulfide)
 - other coloring agents still unknown, complex organics?!

Atmospheric Features

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Zones:

- upwardly moving material
- high, cool cloud bands

Belts:

- falling gasses
- low, warm, and dark

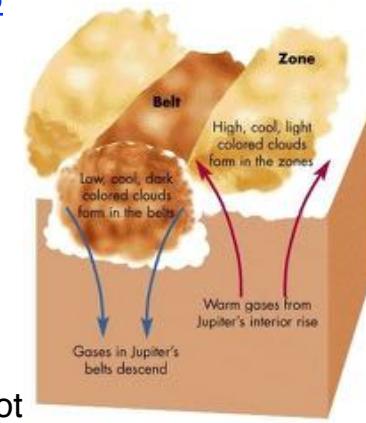
Spots:

- updrafts: white spots
- holes in cloud decks: dark spot

Winds and Circulation: Zonal Winds:

- belts and zones rotate at different rates
- lots of shear from band to zone (+/- 360 km/hr!)
- circulation at interfaces

Example: the Great Red Spot



Atmospheric circulation - belts and zones

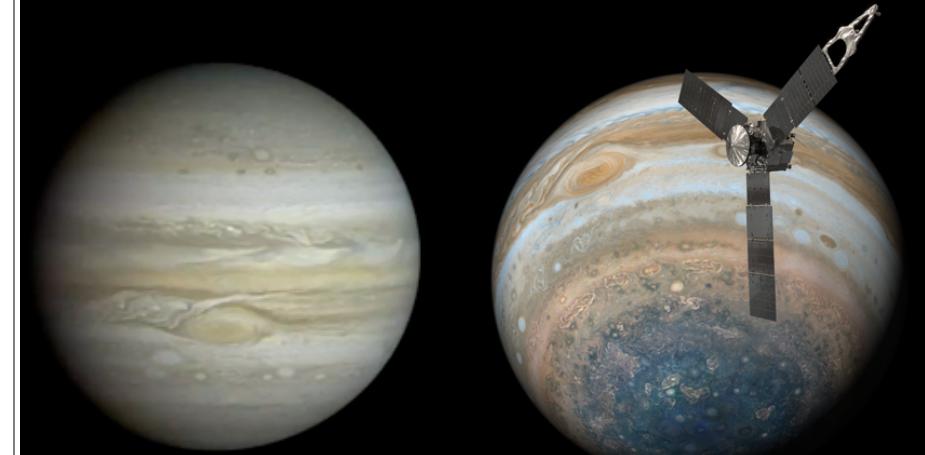


Hubble Space Telescope
January 2015

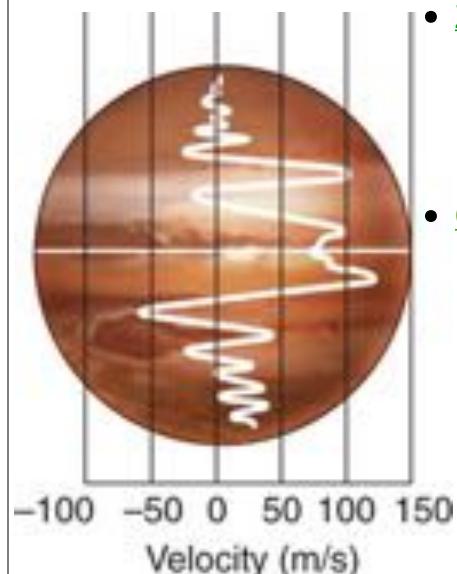


Voyager 1 flyby - 1979
rotation removed

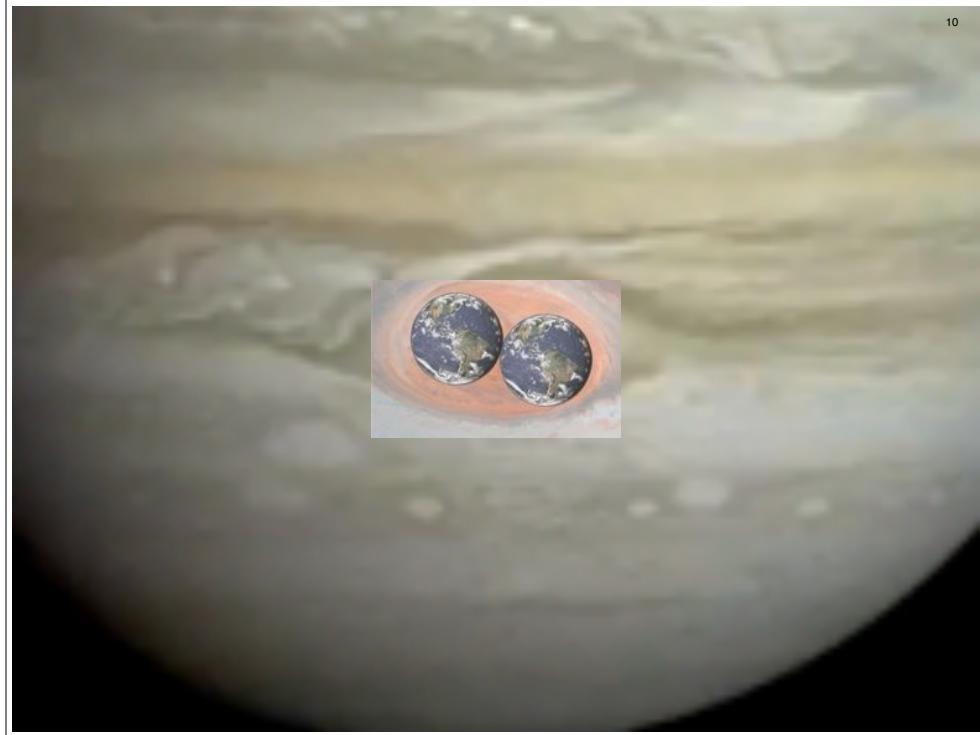
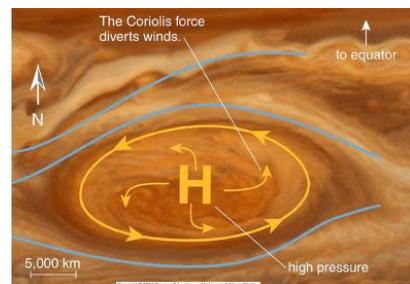
Atmospheric circulation - belts and zones



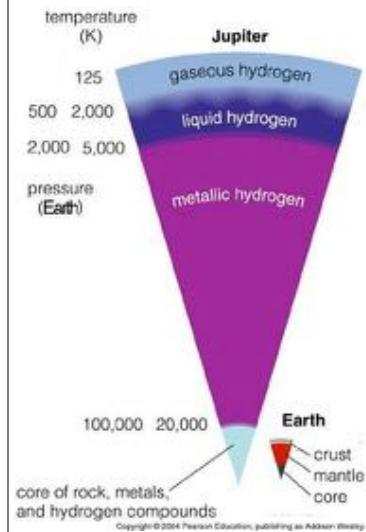
Winds and Circulation



- Zonal Winds:
- belts and zones rotate at different rates
- lots of shear from band to zone (+/-360 km/hr!)
- Circulation at interfaces
Example: the Great Red Spot



The Jovian Interior



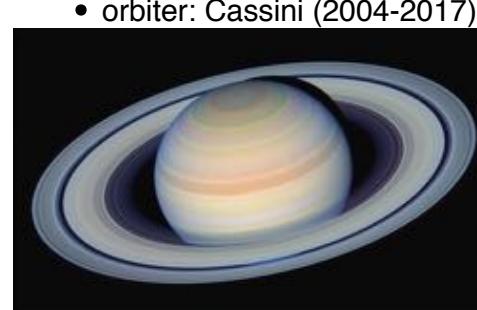
- Clouds and atmosphere
 - temperature increases with depth
 - pressure increases with depth
- Deeper, pressure liquifies hydrogen
 - liquid molecular hydrogen “mantle”
 - metallic liquid hydrogen below (electrons free to roam)
 - starts where $P \sim 2-3000 \times$ Earth’s
 - root of Jovian magnetic field
- A **rocky core** ($M \sim 15 \times$ Earth)
 - $T \sim 20,000 K$, $P \sim 100,000 \times$ Earth

Saturn

- **The Basics:**
 - Mass = 95 x Earth
 - Diameter = 9.5 x Earth
 - “Surface” Gravity = 1.07 x Earth
 - Moons: 5 major, many minor
 - Rotation: very fast: $P \sim 10.6$ hours
 - visits:
 - flybys: Voyager 1, 2 (1981)
 - orbiter: Cassini (2004-2017)

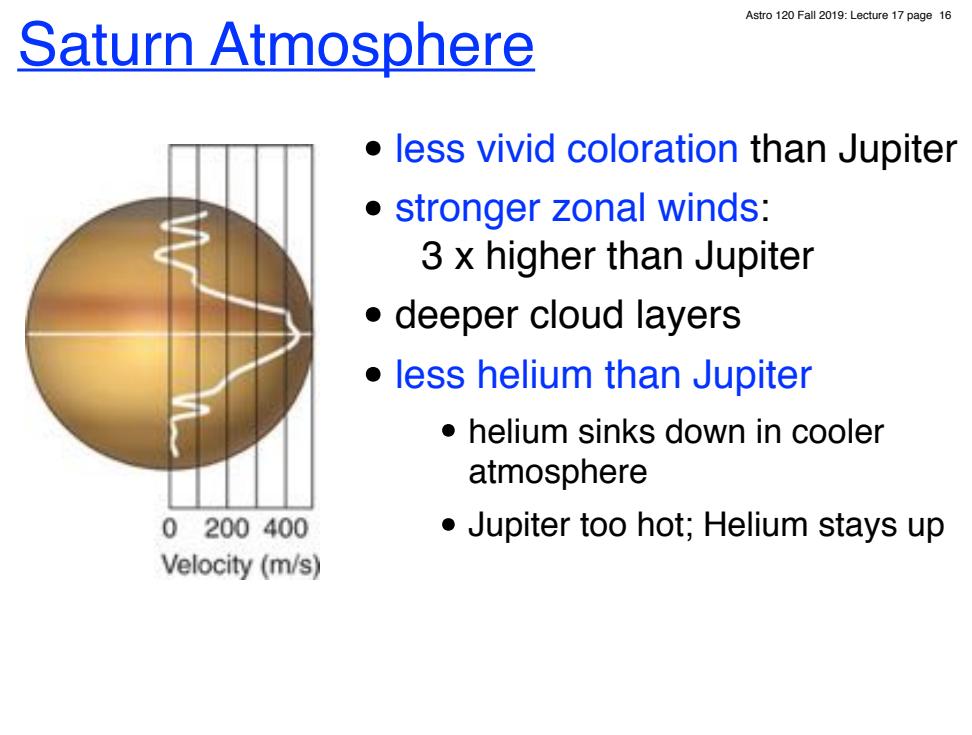
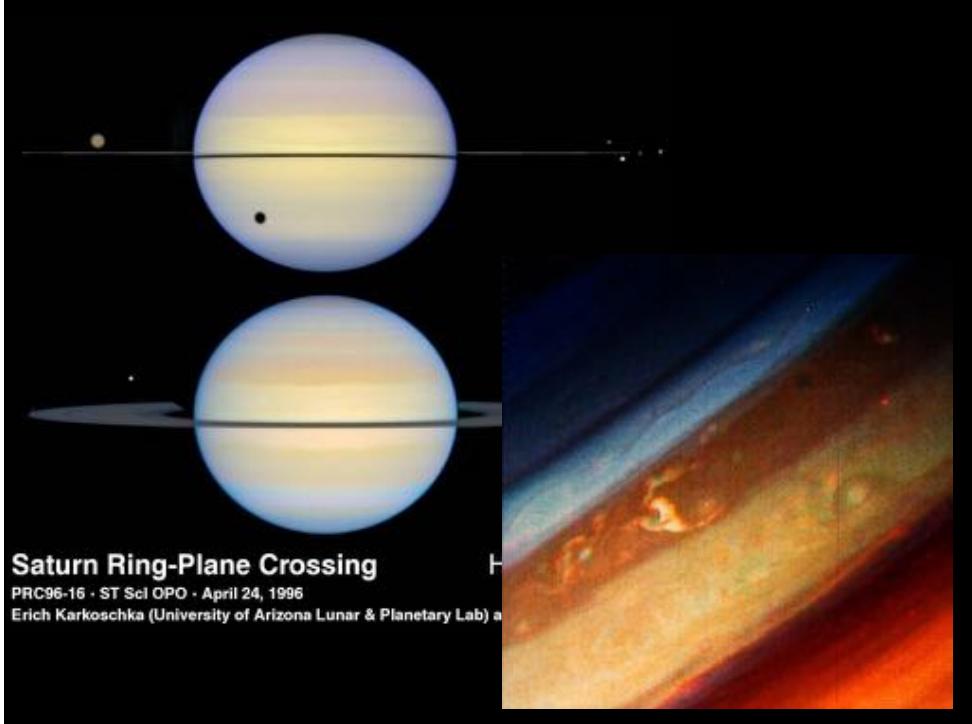


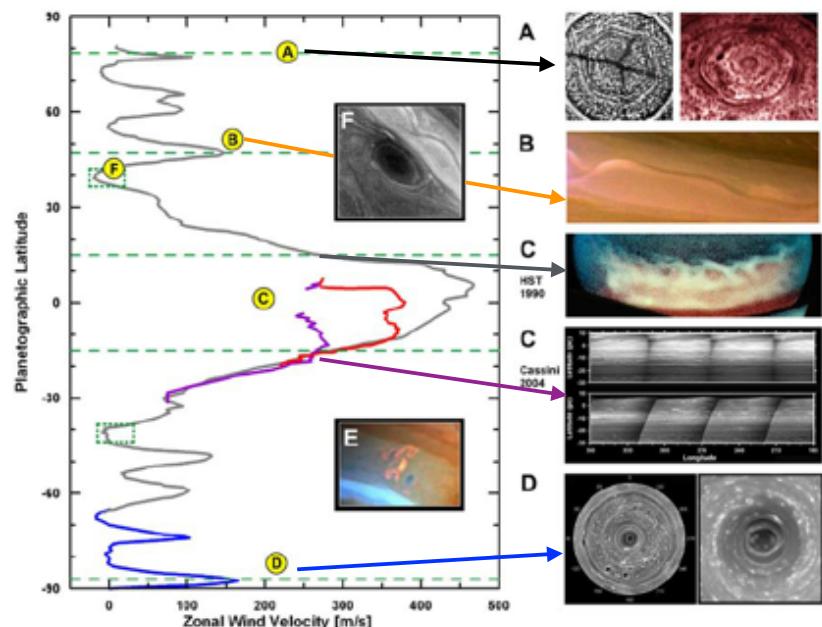
small telescope view



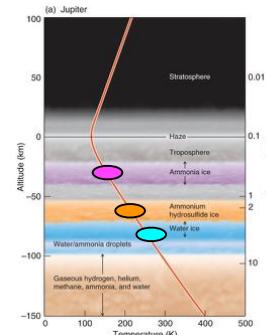
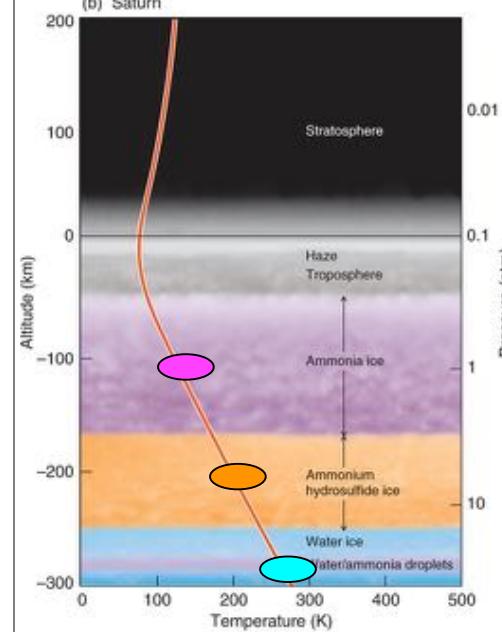
Looking back from Saturn to the Earth...

Saturn





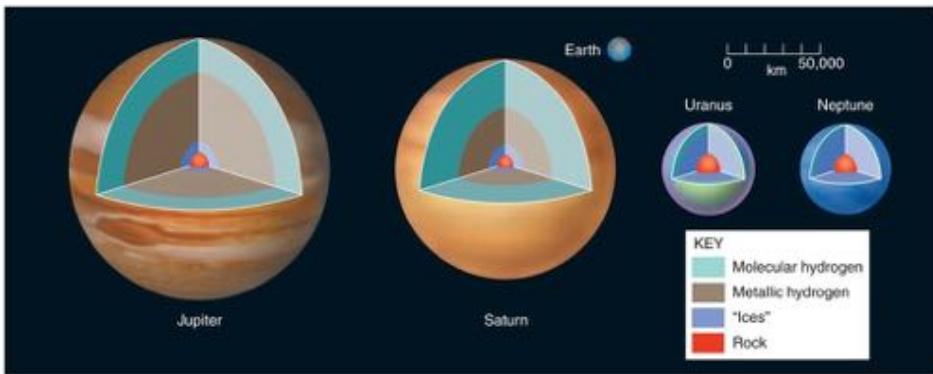
Clouds on Saturn (and Jupiter)



Jupiter's clouds are higher
- warmer close to the sun

Saturn's Interior

- thicker liquid molecular H “mantle”
- smaller metallic H interior
- similar rocky core (to preserve mean density)

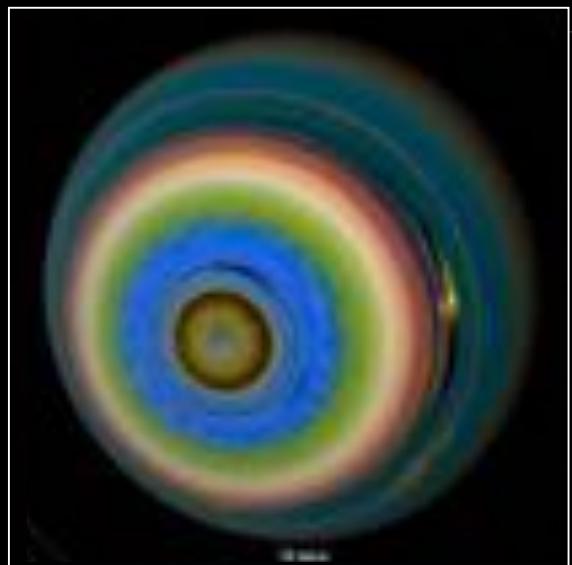


Uranus and Neptune

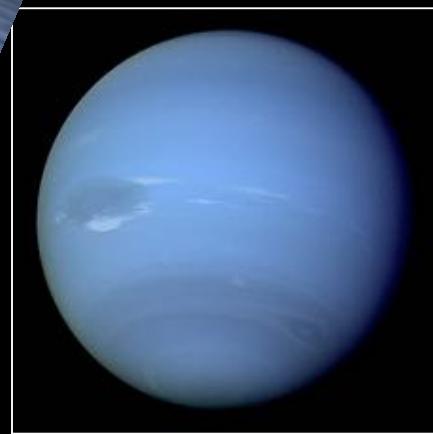
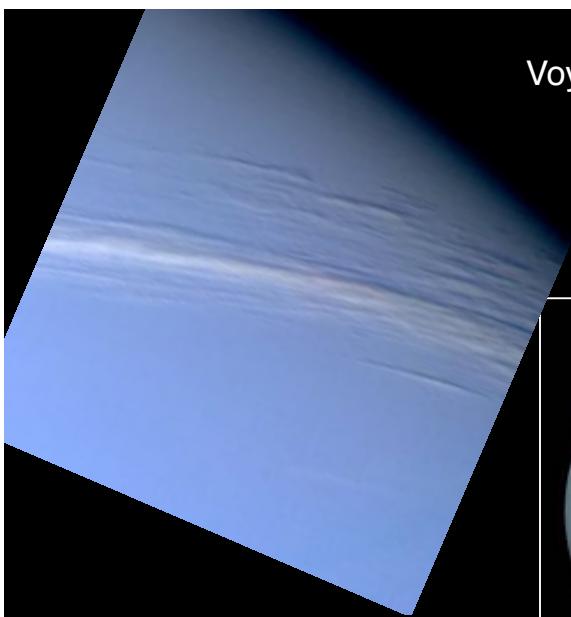
The Basics:
Mass = 14.5 x Earth
Diameter = 4.0 x Earth
Surface Gravity = 0.90 x Earth
Rings: several thin dusky rings
Moons: 5 major, ~10 minor
view from Earth:
• dusky blue disk
• discovered in 1781
visit: flyby: Voyager 2 (1986)
The Basics:
Mass = 17.1 x Earth
Diameter = 3.9 x Earth
Surface Gravity = 1.12 x Earth
Rings: several partial ring arcs
Moons: 1 major, ~7 minor
view from Earth:
• dusky blue disk
• discovered in 1846
visit: flyby: Voyager 2 (1989)



Voyager 2 at Uranus, 1986

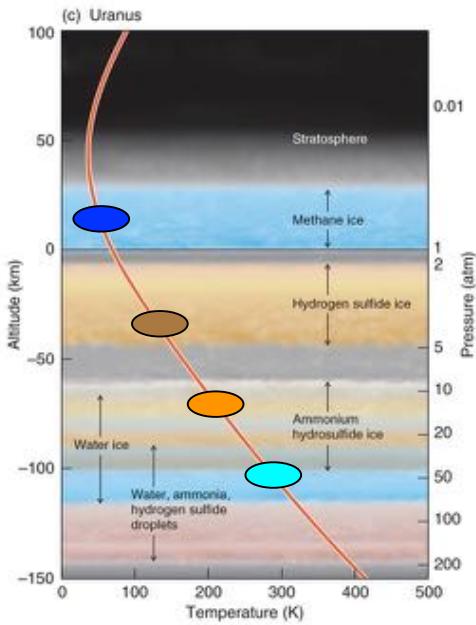


Voyager 2 at Neptune, 1989



Uranus / Neptune Clouds

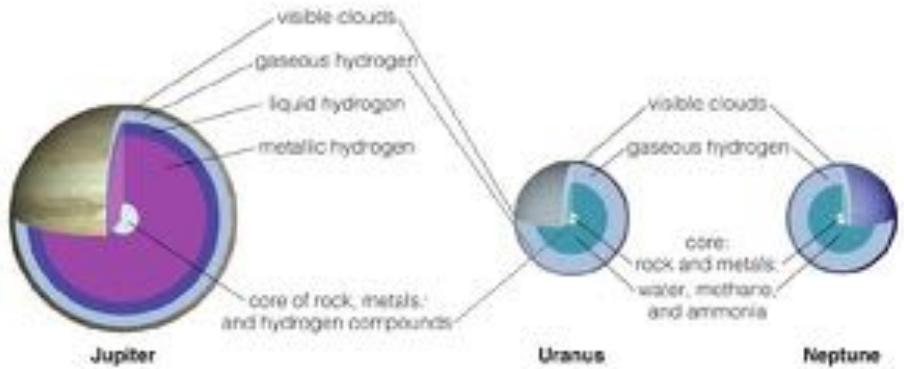
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- Much farther from the Sun
- Colder
- upper atmosphere cold enough for methane to freeze
- blue cloudtops
- deeper, warmer layers:
 - hydrogen sulfide
 - ammonium hydrosulfide
 - water

Possible Interior Structure of Uranus and Neptune

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Outer planets, compared

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