Astro 120 Fall 2019: Lecture 9 page 1

Reading: Chap. 3, Sect. 3.4 - 3.5 Web-based article on orbits for Thursday
Homework 4: Due Thursday for early grading or Mon./Fri.
Exam 1: Next Tuesday, Oct. 1, see schedule on website.
Essay provided in lecture this Thursday

Last time: Kepler to Newton

- Kepler's Laws
 - Empirical: based on data alone w/o physical bias
 - #1: ellipses; #2: planets move faster when close; #3: $P^2 = a^3$
- Newton!
 - gravity as the physical law orbits are continual falls

Today: more Newton

- Physical Laws and definitions of force, velocity, acceleration
- #1: Inertia; #2: Forces (F=ma); #3: Action/Reaction
- Newton's Law of Universal Gravitation
 - gravity as a central, universal, cosmic force



1666: <u>Isaac Newton (1643-1727)</u>

mathematician: Invented calculus as a youth . . .



SYNTHESIZED:

Galileo's Experiments

Kepler's Laws

Calculus into Physical Laws; the basis of <u>Modern Science</u>

Apple falls -> Earth and apple attract each other Moon and Earth attract each other, too

If moon moves sideways as it falls, it could forever circle the Earth...

Newton's Synthesis

Astro 120 Fall 2019: Lecture 9 page 4

- Mathematics Calculus
 - How to define/formulate/calculate motion & acceleration
- Physics definitions / laws
 - energy of interaction between masses
 - momentum resistance to change in motion
 - correspondence with mathematical definitions

• Universal Gravitation

- dependence of gravitational force on mass & distance
- connecting Galileo's experiments & Kepler's Laws
- successful synthesis of earthly & cosmic behavior
- blueprint for modern physics

Newton's Laws: Newton #1: The Law of Inertia

A body moves at a constant velocity unless an unbalanced force acts on it

- <u>Velocity:</u> speed and direction
 - example: 65 mph southbound
- Force: something that changes a body's velocity
 - something that changes body's speed and/or direction
 - an external "push" or "pull"
- Inertia: resistance to change in velocity

Newton's Laws:

Astro 120 Fall 2019: Lecture 9 page 7

mass vs. weight

mass <-> inertia

weight <-> force

Newton #2: The Law of Force Force = mass x acceleration

- Acceleration:
 - (rate of) change in velocity
 - = (rate of) change in speed and/or direction
 - examples:
 - 0mph to 60 mph in 12 seconds (accel.)
 - 60 to 0 in 10 seconds (decel.)
 - turning left at the light (change in direction)
- Inertia: a=F/m
 - bigger mass: smaller accel. for same force
 - inertia: resistance to acceleration by a force
 - example: linebackers are big, wide receivers are small
 - example: shot put vs. golf ball





Astro 120 Fall 2019: Lecture 9 page 5

Newton's Laws:

Newton #3: Law of Action and Reaction

When one body exerts a force on a second body, the second body exerts an equal force, in the opposite direction, on the first.



Newton's Law of Universal Gravitation

Gravity is

- a central force: strength drops with distance²
- a universal force: same form everywhere
- a cosmic force: inherent property of matter

Apple falls -> Earth and apple attract each other Moon and Earth attract each other, too

If moon moves sideways as it falls, it could forever circle the Earth...

Force of gravity pulls planets towards Sun

(Newton's 2nd law)

 without gravity, planets would fly away in straight lines (Newton's 1st law)



Newton's Derivation of Kepler #3

Gravitational force pulling planets toward sun

 $F_{\text{toward}} = \frac{GMm}{a^2}$ (Newton's law of Universal Gravitation)

• centrifugal "force" pulling planets away from sun

$$F_{\text{away}} = rac{mv^2}{a}$$
 or, since $v = rac{2\pi a}{P}$
 $\boxed{F_{\text{away}} = rac{m4\pi^2 a}{P^2}}$

• If forces equal, then distance between doesn't change!

$$\frac{GMm}{a^2} = \frac{m4\pi^2 a}{P^2}$$
 ... or ... $P^2 = a^3 \times \left(\frac{4\pi^2}{GM}\right)^2$ a constant

this is Kepler's Third Law!

Newton's Legacy

• Force of Gravity pulls planets towards Sun

Astro 120 Fall 2019: Lecture 9 page 13

- without gravity, planets would fly away in straight lines
- Newton's theory of gravity explains -simply- the orbits of the planets

Understanding motions of the planets was the principal discovery of astronomy from prehistory through 1700.

- Improved observations ("technology") demanded more precise models of the Solar System
- This precision was
 - <u>approached</u> by complex models (epicycles, etc.) but
 - achieved by discovery of the underlying simplicity: Gravity

